

# GCSE Astronomy      Topic 2 Key Facts

## The Solar System

Our solar system contains 8 **planets** –Mercury, Venus, Earth, Jupiter, Saturn, Uranus and Neptune

There are also 5 **dwarf planets** – Ceres (asteroid belt), Pluto, Haumea and Makemake (Kuiper belt), Eris (scattered disc)

**Asteroids** are rocks greater than 10m in diameter. Most are in a belt between Mars and Jupiter. They have no coma.

**Comets** have a coma (fuzzy atmosphere). They are balls of rock and ice. Their orbits are very eccentric and often retrograde.

**Centaur**s are a cross between asteroids and comets. Their orbits are eccentric but they stay inside the solar system. Some have a coma. They include Hidalgo and Chiron.

**Trans Neptunian Objects** (TNO's) orbit beyond Neptune.

**The Kuiper Belt** is beyond Neptune. It is a belt of millions of small objects between 30 and 50 AU and is possibly the origin of short period comets. The solar system formed from a disc. The inner part formed the planets and the outer part the Kuiper belt. When the Kuiper belt thins it becomes known as the Scattered Disc.

**The Oort cloud** is a theoretical huge cloud (50,000 AU) surrounding the solar system. It is possibly the source of long period comets. It could be the remnants of the early solar system, i.e. the planetary nebula from which it formed.

## The Planets in Orbit

The Earth orbits the Sun on the ecliptical plane. The orbit of other planets is slightly inclined to the ecliptic e.g. the inclination of Mercury is greatest at 7 degrees. The average distance to the Sun is 1 AU.

Planet comes from a word meaning wandering star. They appear to move across the sky in a band called the zodiac.

Sometimes planets appear to move backwards. This is called retrograde motion.

- **Perihelion** – when a planet is closest to the Sun (e.g. the Earth in January)
- **Aphelion** – when a planets is furthest from the Sun (e.g. the Earth in December)
- **Greatest elongation** – the greatest angular separation of an inferior planet from the Sun (e.g. Venus is 45 degrees)
- **Conjunction** – when 2 or more planets appear close in the sky
- **Opposition** – when a superior planet is on the opposite side of the Earth to the Sun (so is best viewed)
- **Transit** – when a smaller body passes in front of a larger one (e.g. the transit of Venus across the Sun)
- **Occultation** – when a body is hidden by another (e.g. when a planet passes behind the Sun)

## The Planets

The first 4 planets are rocky. The next 4 are gas giants.

**Mercury** – surface like our moon's. No atmosphere. Revolves slowly so big variation in temperature.

**Venus** – similar size to Earth. Very dense CO<sub>2</sub> atmosphere makes the surface hotter than Mercury (acts like a blanket to keep the heat in). Show phases viewed from earth. Radar has been used to map its surface.

**Mars** – Appears red in colour. Canyons and valleys are evidence of erosion. Very thin atmosphere mostly CO<sub>2</sub>.

**Jupiter** – The largest planet. High velocity winds on the surface and a huge red spot. 90% hydrogen with a core of metallic hydrogen.

**Saturn** – another gas giant. Quite squashed due to its high rotational speed.

Saturn has a distinctive ring system about 1 km thick made up of particles from 1cm to 1m in size. There are 3 main regions, A and B separated by the Cassini Division then C.

**Uranus** – spins on a very tilted axis so its south pole almost faces the Sun. Blue colour due to some methane. Was discovered by William Herschel with his powerful new telescope.

**Neptune** – the last gas giant. Similar composition to Uranus. Its existence was predicted due to the effect of its gravity on Uranus.

## The origin of Moons

Our moon is probably formed from the debris left by a massive impact between the Earth and another body

Mars has 2 small moons, Phobos and Demos, which are probably captured asteroids

Neptune's moon Triton has a retrograde orbit. This suggests that it formed at a later time and is probably a trapped Kuiper belt object.

## Comets

Comets have an icy rocky nucleus and a coma (visible fuzzy atmosphere).

They have two tails which grow greatly in size as the comet approaches the Sun and its surface is blown away.

An ion tail which always points away from the Sun due to the solar wind

A dust tail which is curved as the particles in it travel slower because they are further from the Sun

Comet orbits are very eccentric, often retrograde and very inclined to the ecliptic. This suggests that their origin is outside the main solar system.

## Meteors

- Meteoroid – a ball of rock in space smaller than an asteroid
- Meteor – one that has entered the atmosphere and is burning up leaving a fiery tail
- Fireball – a very bright meteor
- Meteorite – a meteor that lands on Earth
- Micrometeorites – tiny meteorites, tons of which fall on earth every day

As comets orbit the Sun they leave a trail of debris. The Earth passes through some of these at the same time every year resulting in yearly meteor showers, e.g. the Leonids appear in Leo every year.

A typical shower may have one meteor per minute. If we trace their paths back they all originate from the same point called the **radiant**.

## Potentially Hazardous Objects (P.H.O.s)

To be potentially dangerous an object must be at least 150m in diameter and pass within 0.05 AU of Earth.

There is evidence that such objects have hit earth in the past, e.g. a giant crater in Arizona.

NASA have, at present, identified just over 1000 objects like these and are keeping a close eye on them. If one hit the sea it could cause a massive tsunami.

NASA has observed the impacts of such bodies with Jupiter

## Famous Discoveries

Before **Copernicus** people believed the Earth was the centre of the Universe. Copernicus suggested a Sun centred model which was simpler based on the available evidence

**Tycho** took many accurate observations of the motions of stars and planets

**Kepler's** three laws of planetary motion (formulated mostly using Tycho's data)

- 1<sup>st</sup> law - The orbit of a planet / comet about the Sun is an ellipse with the Sun at one focus
- 2<sup>nd</sup> law - A line joining a planet to the Sun sweeps out equal areas in equal time intervals
- 3<sup>rd</sup> law - The squares of the periods of the planets are proportional to the cubes of their mean distance from the Sun. If T is in years and distance in AU then  $T^2 = R^3$

**Galileo** built his own telescope and made many important discoveries including ...

- Venus showed phases like the Moon (evidence that it orbits the Sun)
- Several moons orbit Jupiter now known as the Galilean moons (Callisto, Europa, Ganymede, Io)
- The surface of the Moon shows mountains and other features

## Other Discoveries

**Ceres** – the first asteroid discovered by Giuseppe Piazzi

**Uranus** – the first planet discovered using a telescope by William Herschel

**Neptune** – its existence was predicted from irregularities in the orbit of Uranus. It was then found by Heinrich D'Arrest

**Pluto** – its existence was wrongly predicted but then it was found by accident by Clyde Tombaugh using a photographic technique called a blink comparator

## Gravity

All masses attract each other. The attractive force gets weaker with distance.

If the distance is x 2 the force is  $\frac{1}{4}$ . If the distance is x 3 the force is  $\frac{1}{9}$ . We call this an inverse square law.

Smaller bodies orbit larger bodies due to the force of gravity. **Newton** first realised that it was the Earth's gravity pulling on the Moon that kept it in orbit around the Earth.

## Exoplanets

These are planets which orbit stars other than the Sun. They are too far to be seen directly (all stars are just points of light) but there are ways of detecting them indirectly

- **Astrometry** – very accurate measurement of the position of stars may reveal “wobbles” due to the gravitational influence of large planets
- **Doppler shifts** – may also reveal stars wobbling due to the gravity of large planets
- **Transits** – the light from a star may dim periodically as a large planet crosses in front of it

## Life on Other Planets

Life (as we know it) requires liquid water. It may have originated on Earth in several ways

- Comets and meteorites brought water to the prehistoric Earth (the Rosetta probe will study this)
- Numerous chemical reactions may have produced enough water to form oceans
- Bacteria may have produced water

**Habitable Zones** are otherwise known as Goldilocks zones, they are regions around a star where the temperature is such that life could evolve on a planet in the zone

**The Drake Equation** calculates the number of civilisations present that we may be able to communicate with. To do this it estimates \* how many stars are formed \* how many have planets \* how many could support life \* on how many does life evolve \* is this life intelligent \* does it develop technology \* how long it survives

## Looking for Life

- Scientists listen for signals from space (The SETI program)
- Probes sent to planets examine samples for signs of past or present life, e.g. fossils

If we discover / make contact with other life forms

- We may learn more about our evolution and place in the Universe, we may be able to share technologies
- We must be careful not to spread harmful diseases, we should hope they are friendly